

# Extreme value analysis and user defined probability functions in Stan

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This notebook demonstrates how to implement user defined probability functions in Stan language. As an example I use generalized Pareto distribution (GPD) and geomagnetic storm data by World Data Center for Geomagnetism.

Load some libraries:

```
library(ggplot2)
library(tidyr)
library(dplyr)
library(gridExtra)
library(rstanarm)
library(rstan)
library(bayesplot)
library(loo)
rstan_options(auto_write = TRUE)
options(mc.cores = parallel::detectCores())
source("stan_utility.R")
```

Read the data. This file has magnitudes of 373 geomagnetic storms which lasted longer than 48h with absolute magnitude larger than 100 in 1957-2014.

```
# file preview shows a header row
d <- read.csv("geomagnetic_tail_data.csv", header = FALSE)
colnames(d) <- "dst"
d <- d %>% mutate(dst = abs(dst)) %>% arrange(dst)
```

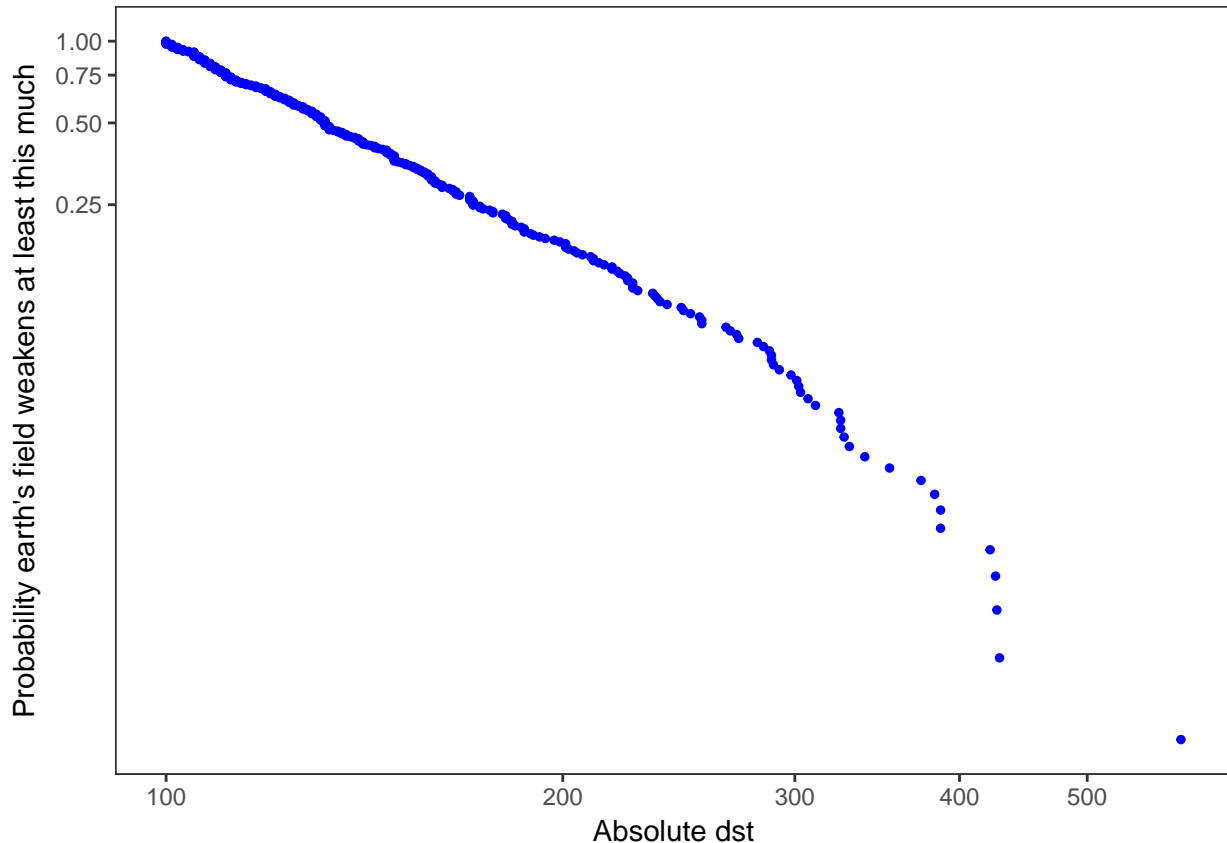
Compute empirical complementary cumulative distribution function.

```
n <- dim(d)[1]
d$ccdf <- seq(n,1,-1)/n
head(d)
```

```
##  dst      ccdf
## 1 100 1.0000000
## 2 100 0.9973190
## 3 100 0.9946381
## 4 100 0.9919571
## 5 100 0.9892761
## 6 100 0.9865952
```

Plot just the data and empirical ccdf

```
ggplot() +
  geom_point(aes(dst, ccdf), data = d, size = 1, colour = "blue") +
  coord_trans(x="log10", y="log10") +
  labs(y = "Probability earth's field weakens at least this much", x = "Absolute dst") +
  guides(linetype = F) +
  theme_bw()
```



The largest event in the data is the March 1989 geomagnetic storm also known as Quebec blackout event ([https://en.wikipedia.org/wiki/March\\_1989\\_geomagnetic\\_storm](https://en.wikipedia.org/wiki/March_1989_geomagnetic_storm)). Now we are interested in estimating the probability of observing a same magnitude event in the future which would be helpful, for example, for insurance companies (for a short term geomagnetic storm predictions there are very elaborate models using observations closer to sun, too). Extreme value theory says that many distributions have a tail which is well modelled with generalized Pareto distribution given the cutoff point is far enough in the tail. For a more detailed model we could also take into account that geomagnetic storms are temporally correlated and tend to appear in bursts.

Generalized Pareto distribution is defined as

$$p(y|u, \sigma, k) = \begin{cases} \frac{1}{\sigma} \left(1 + k \left(\frac{y-u}{\sigma}\right)\right)^{-\frac{1}{k}-1}, & k \neq 0 \\ \frac{1}{\sigma} \exp\left(-\frac{y-u}{\sigma}\right), & k = 0, \end{cases}$$

where  $u$  is a lower bound parameter,  $y$  is restricted to the range  $(u, \infty)$ ,  $\sigma$  is a scale parameter, and  $k$  is a shape parameter (see, e.g., [https://en.wikipedia.org/wiki/Generalized\\_Pareto\\_distribution](https://en.wikipedia.org/wiki/Generalized_Pareto_distribution) for cdf and random number generation).

For probability distributions implemented in Stan math library there are functions

- `_lpdf` (or `_lpmf`): log probability density (mass) function
- `_cdf`: cumulative distribution function
- `_lcdf`: log cumulative distribution function
- `_lccdf`: log complementary cumulative distribution function
- `_rng`: generate random variables from the distribution

The user defined functions can have only one function signature unlike the functions implemented in C++ in Stan math library. In the example I have chosen the function signatures so that the behavior is as close as

possible to most usual ways to call these functions. The main feature is that all these function return a scalar real value (except `_rng` would return scalar integer for a discrete integer valued distribution).

For generalized Pareto distribution we implement

- real `gpareto_lpdf`(vector  $y$  | real  $ymin$ , real  $k$ , real  $sigma$ )
- real `gpareto_cdf`(vector  $y$  | real  $ymin$ , real  $k$ , real  $sigma$ )
- real `gpareto_lcdf`(vector  $y$  | real  $ymin$ , real  $k$ , real  $sigma$ )
- real `gpareto_lccdf`(vector  $y$  | real  $ymin$ , real  $k$ , real  $sigma$ )
- real `gpareto_rng`(real  $ymin$ , real  $k$ , real  $sigma$ )

As we can define only one function signature for user defined functions, I have now chosen to have vector type for  $y$  and real types for the parameters, while builtin functions can handle vectors, arrays and scalars (denoted by generic type `reals`). For vector valued  $y$  `_lpdf`, `_lcdf`, and `_lccdf` return sum of log values computed with each element of  $y$ . For vector valued  $y$  `_cdf` returns product of values computed with each element of  $y$ . `_rng` returns a single random number generated from the generalized Pareto distribution.

The whole code for functions, the basic model and generated quantities for posterior predictive checking (ppc), leave-one-out cross-validation (loo), and prediction of rare events is shown below. I will next go through some practical details.

```
writeLines(readLines("gpareto.stan"))
```

```
## functions {
##   real gpareto_lpdf(vector y, real ymin, real k, real sigma) {
##     // generalised Pareto log pdf
##     int N;
##     N = dims(y)[1];
##     if (k<0 && max(y-ymin)/sigma > -1/k)
##       reject("k<0 and max(y-ymin)/sigma > -1/k; found k, sigma =", k, sigma)
##     if (sigma<=0)
##       reject("sigma<=0; found sigma =", sigma)
##     if (fabs(k) > 1e-15)
##       return -(1+1/k)*sum(log1p((y-ymin) * (k/sigma))) -N*log(sigma);
##     else
##       return -sum(y-ymin)/sigma -N*log(sigma); // limit k->0
##   }
##   real gpareto_cdf(vector y, real ymin, real k, real sigma) {
##     // generalised Pareto cdf
##     if (k<0 && max(y-ymin)/sigma > -1/k)
##       reject("k<0 and max(y-ymin)/sigma > -1/k; found k, sigma =", k, sigma)
##     if (sigma<=0)
##       reject("sigma<=0; found sigma =", sigma)
##     if (fabs(k) > 1e-15)
##       return exp(sum(log1m_exp((-1/k)*(log1p((y-ymin) * (k/sigma))))));
##     else
##       return exp(sum(log1m_exp(-(y-ymin)/sigma))); // limit k->0
##   }
##   real gpareto_lcdf(vector y, real ymin, real k, real sigma) {
##     // generalised Pareto log cdf
##     if (k<0 && max(y-ymin)/sigma > -1/k)
##       reject("k<0 and max(y-ymin)/sigma > -1/k; found k, sigma =", k, sigma)
##     if (sigma<=0)
##       reject("sigma<=0; found sigma =", sigma)
##     if (fabs(k) > 1e-15)
##       return sum(log1m_exp((-1/k)*(log1p((y-ymin) * (k/sigma)))));
##     else
```

```

##     return sum(log1m_exp(-(y-ymin)/sigma)); // limit k->0
##   }
##   real gpareto_lccdf(vector y, real ymin, real k, real sigma) {
##     // generalised Pareto log ccdf
##     if (k<0 && max(y-ymin)/sigma > -1/k)
##       reject("k<0 and max(y-ymin)/sigma > -1/k; found k, sigma =", k, sigma)
##     if (sigma<=0)
##       reject("sigma<=0; found sigma =", sigma)
##     if (fabs(k) > 1e-15)
##       return (-1/k)*sum(log1p((y-ymin) * (k/sigma)));
##     else
##       return -sum(y-ymin)/sigma; // limit k->0
##   }
##   real gpareto_rng(real ymin, real k, real sigma) {
##     // generalised Pareto rng
##     if (sigma<=0)
##       reject("sigma<=0; found sigma =", sigma)
##     if (fabs(k) > 1e-15)
##       return (ymin + (pow(uniform_rng(0,1),-k)-1)*sigma/k);
##     else
##       return (ymin - sigma*log(1/uniform_rng(0,1)-1)); // limit k->0
##   }
## }
## data {
##   real ymin;
##   int<lower=0> N;
##   vector<lower=ymin>[N] y;
##   int<lower=0> Nt;
##   vector<lower=ymin>[Nt] yt;
## }
## transformed data {
##   real ymax;
##   ymax = max(y);
## }
## parameters {
##   real<lower=0> sigma;
##   real<lower=-sigma/(ymax-ymin)> k;
## }
## model {
##   y ~ gpareto(ymin, k, sigma);
## }
## generated quantities {
##   vector[N] log_lik;
##   vector[N] yrep;
##   vector[Nt] predccdf;
##   for (n in 1:N) {
##     log_lik[n] = gpareto_lpdf(rep_vector(y[n],1) | ymin, k, sigma);
##     yrep[n] = gpareto_rng(ymin, k, sigma);
##   }
##   for (nt in 1:Nt)
##     predccdf[nt] = exp(gpareto_lccdf(rep_vector(yt[nt],1) | ymin, k, sigma));
## }

```

For each function we do basic argument checking. In addition of invalid values due to user errors, due to the

limited accuracy of the floating point presentation of the values, sometimes we may get invalid parameters, for example. `sigma` equal to zero, even if we have declared it as `real<lower=0> sigma;`. For these latter cases, it is useful to use `reject` statement, and the corresponding MCMC proposal will be rejected and the sampling can continue.

```

    if (k<0 && max(y-ymin)/sigma > -1/k)
      reject("k<0 and max(y-ymin)/sigma > -1/k; found k, sigma =", k, sigma)
    if (sigma<=0)
      reject("sigma<=0; found sigma =", sigma)

```

Stan documentation warns about use of `fabs` and conditional evaluation as they may lead to discontinuous energy or gradient. In GPD we need to handle a special case of  $k \rightarrow 0$ . The following will cause discontinuity, but the magnitude of the discontinuity is the same order as the accuracy of the floating point presentation, and thus we are not

```

    if (fabs(k) > 1e-15)
      return -(1+1/k)*sum(log1p((y-ymin) * (k/sigma))) -N*log(sigma);
    else
      return -sum(y-ymin)/sigma -N*log(sigma); // limit k->0

```

In data block we need to define the minimum threshold `ymin`. We also define test points `yt`.

```

data {
  real ymin;
  int<lower=0> N;
  vector<lower=ymin>[N] y;
  int<lower=0> Nt;
  vector<lower=ymin>[Nt] yt;
}

```

`sigma` has to be positive and `k` has a lower limit which depends on `sigma` and the maximum value.

```

transformed data {
  real ymax;
  ymax = max(y);
}
parameters {
  real<lower=0> sigma;
  real<lower=-sigma/(ymax-ymin)> k;
}

```

By defining `gpareto_lpdf` we can use also the common `~` notation in Stan to write familiar looking model code.

```

model {
  y ~ gpareto(ymin, k, sigma);
}

```

In generated quantities we compute `log_lik` values for LOO and `yrep` values for posterior predictive checking. We had defined the first argument of `pareto_lpdf` to be a vector. Now a single element of vector `y[n]` has a type `real`, and as Stan has strong typing and we can't overload the user defined functions, we need to cast `y[n]` to be a vector by making a 1 element vector `rep_vector(y[n],1)`. Alternatively we could write another function with a different name which could accept scalar `y[n]`, but here I wanted to demonstrate the minimal approach.

```

for (n in 1:N) {
  log_lik[n] = gpareto_lpdf(rep_vector(y[n],1) | ymin, k, sigma);
  yrep[n] = gpareto_rng(ymin, k, sigma);
}

```

Finally we compute predictive probabilities for observing events with certain magnitudes.

```
for (nt in 1:Nt)
  predccdf[nt] = exp(gpareto_lccdf(rep_vector(yt[nt],1) | ymin, k, sigma));
```

Next we use the defined Stan model to analyse the distribution of the largest geomagnetic storms.

Fit the Stan model

```
yt<-append(10^seq(2,3,.01),850)
ds<-list(ymin=100, N=n, y=d$dst, Nt=length(yt), yt=yt)
fit_gpd <- stan(file='gpareto.stan', data=ds, refresh=0,
               chains=4, seed=100) #850
```

```
## trying deprecated constructor; please alert package maintainer
```

Run the usual diagnostics

```
check_treedepth(fit_gpd)
```

```
## [1] "0 of 4000 iterations saturated the maximum tree depth of 10 (0%)"
```

```
check_energy(fit_gpd)
```

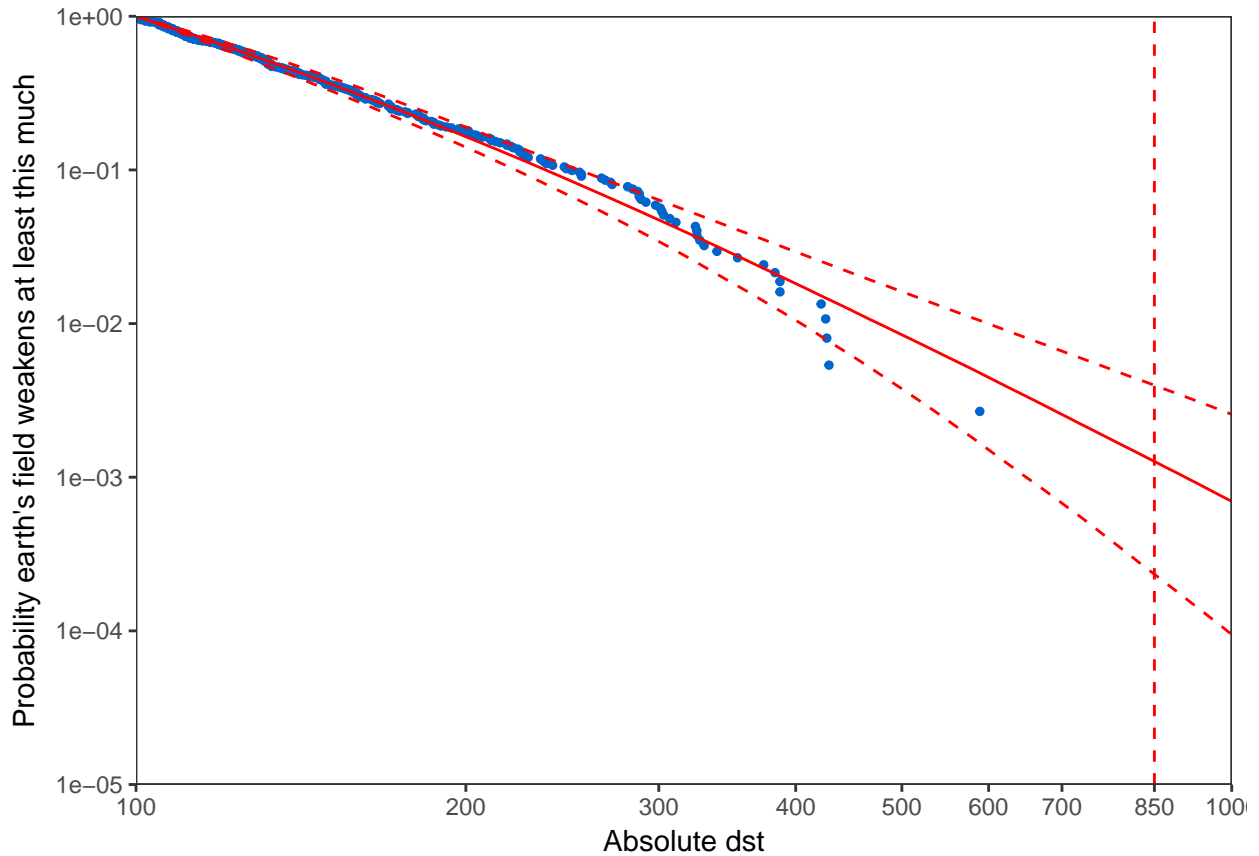
```
check_div(fit_gpd)
```

```
## [1] "0 of 4000 iterations ended with a divergence (0%)"
```

Diagnostics do not find anything alarming.

Plot the model fit with 90% posterior interval. The largest observed magnitude in the data corresponds to Quebec blackout in 1989 ([https://en.wikipedia.org/wiki/March\\_1989\\_geomagnetic\\_storm](https://en.wikipedia.org/wiki/March_1989_geomagnetic_storm)). The vertical dashed line at 850 shows the estimated magnitude of Carrington event in 1859 ([https://en.wikipedia.org/wiki/Solar\\_storm\\_of\\_1859](https://en.wikipedia.org/wiki/Solar_storm_of_1859)).

```
gpd_params <- rstan::extract(fit_gpd)
mu <- apply(t(gpd_params$predccdf), 1, quantile, c(0.05, 0.5, 0.95)) %>%
  t() %>% data.frame(x = yt, .) %>% gather(pct, y, -x)
clrs <- color_scheme_get("brightblue")
ggplot() +
  geom_point(aes(dst, ccdf), data = d, size = 1, color = clrs[[5]]) +
  geom_line(aes(x=c(850,850),y=c(1e-5,1)),linetype="dashed",color="red") +
  geom_line(aes(x, y, linetype = pct), data = mu, color = 'red') +
  scale_linetype_manual(values = c(2,1,2)) +
  coord_trans(x="log10", y="log10", limx=c(100,1000), limy=c(1e-5,1)) +
  scale_y_continuous(breaks=c(1e-5,1e-4,1e-3,1e-2,1e-1,1), limits=c(1e-5,1)) +
  scale_x_continuous(breaks=c(100,200,300,400,500,600,700,850,1000), limits=c(100,1000)) +
  labs(y = "Probability earth's field weakens at least this much", x = "Absolute dst") +
  guides(linetype = F) +
  theme_bw()
```



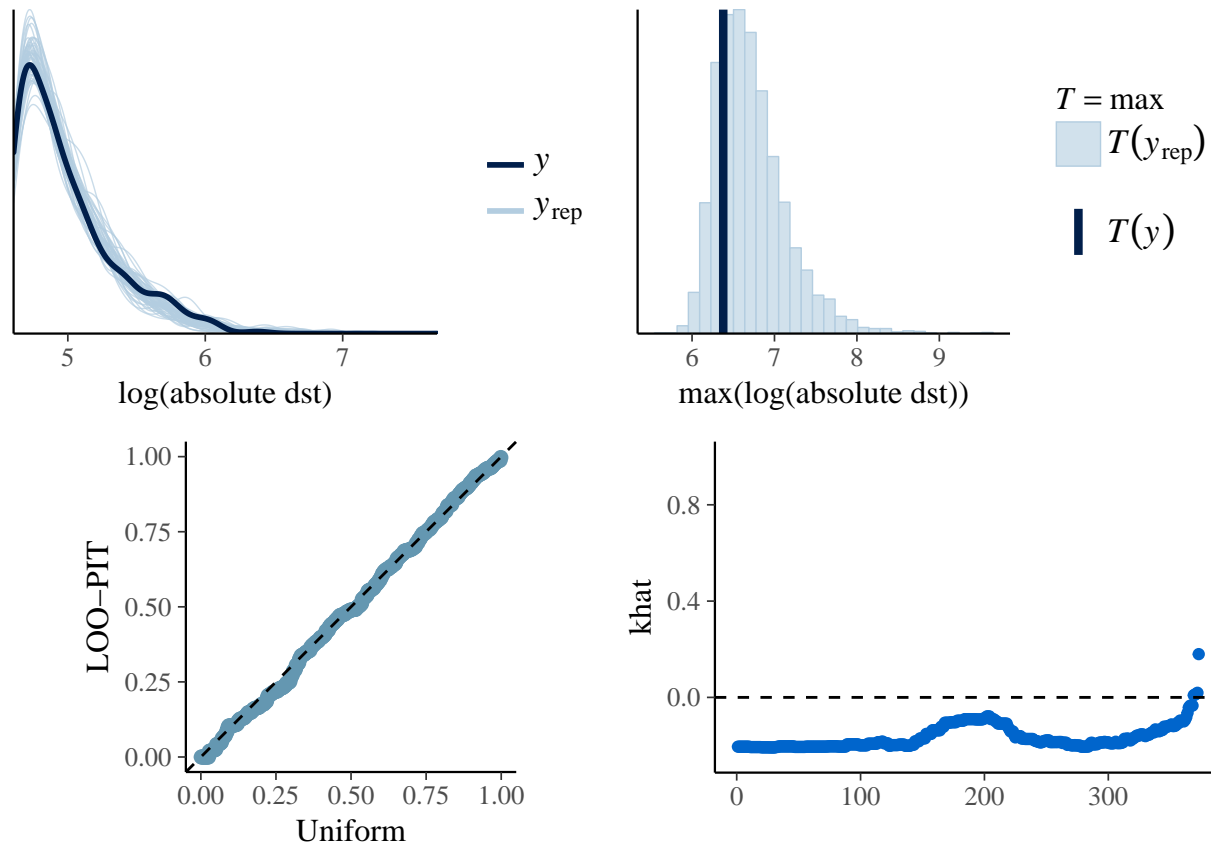
For additional model checking plot 1) kernel density estimate of  $\log(\text{dst})$  and posterior predictive replicates, 2) max log magnitude (Quebec event) and histogram of maximums of posterior predictive replicates, 3) leave-one-out cross-validation probability-integral-transformation, and 4)  $\hat{\kappa}$  values from the PSIS-LOO. None of these diagnostics can find problems in the model fit.

```

ppc1 <- ppc_dens_overlay(log(d$dst), log(gpd_params$yrep[1:50,])) + labs(x="log(absolute dst)")
ppc2 <- ppc_stat(log(d$dst), log(gpd_params$yrep), stat = "max") + labs(x="max(log(absolute dst))")
psis <- psislw(-gpd_params$log_lik)
clrs <- color_scheme_get("brightblue")
pkhats <- ggplot() + geom_point(aes(x=seq(1,n),y=psis$pareto_k), color=clrs[[5]]) + labs(y="khat", x="")
  geom_hline(yintercept=0, linetype="dashed") + ylim(-0.25,1) + theme_default()
ppc3 <- ppc_loo_pit(log(d$dst), log(gpd_params$yrep), lw=psis$lw_smooth)
grid.arrange(ppc1,ppc2,ppc3,pkhats,ncol=2)

```

## `stat_bin()` using `bins = 30`. Pick better value with `binwidth`.



We compute also PSIS-LOO estimate, although this is not much useful without alternative model to compare.

```
(loo_gpd<-loo(gpd_params$log_lik))
```

```
## Computed from 4000 by 373 log-likelihood matrix
##
##      Estimate  SE
## elpd_loo -1874.7 23.7
## p_loo      1.8  0.2
## looic      3749.4 47.3
##
## All Pareto k estimates are good (k < 0.5)
## See help('pareto-k-diagnostic') for details.
```

Here the analysis was simplified by assuming the geomagnetic storm events to be independent in time although they appear in bursts. Generalized Pareto distribution is valid fit also for correlated data, but for predicting the probability of Quebec blackout magnitude event in the future we should take into account also the correlation. Based on the simplified assumption the probability that we would observe Quebec blackout magnitude event in the next 57 years is 80% and the probability of observing Carrington level event in the next 57 years is 40%. Now go read the effects of these geomagnetic storms.

## Appendix: Session information

```
sessionInfo()
```

```
## R version 3.2.3 (2015-12-10)
## Platform: x86_64-pc-linux-gnu (64-bit)
```



```

## Running under: Ubuntu 16.04.3 LTS
##
## locale:
## [1] LC_CTYPE=en_US.UTF-8      LC_NUMERIC=C
## [3] LC_TIME=en_US.utf8        LC_COLLATE=en_US.UTF-8
## [5] LC_MONETARY=en_US.UTF-8   LC_MESSAGES=en_US.UTF-8
## [7] LC_PAPER=fi_FI.utf8       LC_NAME=C
## [9] LC_ADDRESS=C              LC_TELEPHONE=C
## [11] LC_MEASUREMENT=en_US.UTF-8 LC_IDENTIFICATION=C
##
## attached base packages:
## [1] stats      graphics  grDevices  utils      datasets  methods    base
##
## other attached packages:
## [1] bindrcpp_0.2      loo_1.1.0      bayesplot_1.4.0
## [4] rstan_2.16.2      StanHeaders_2.16.0-1 rstanarm_2.15.3
## [7] Rcpp_0.12.13      gridExtra_2.3  dplyr_0.7.4
## [10] tidyr_0.7.2       ggplot2_2.2.1
##
## loaded via a namespace (and not attached):
## [1] lattice_0.20-35    zoo_1.8-0      gtools_3.5.0
## [4] assertthat_0.2.0  rprojroot_1.2  digest_0.6.12
## [7] mime_0.5           R6_2.2.2       plyr_1.8.4
## [10] backports_1.1.1   stats4_3.2.3   evaluate_0.10.1
## [13] colourpicker_1.0  rlang_0.1.4    lazyeval_0.2.1
## [16] minqa_1.2.4       miniUI_0.1.1   nloptr_1.0.4
## [19] Matrix_1.2-11     DT_0.2         rmarkdown_1.6
## [22] labeling_0.3      shinythemes_1.1.1 shinyjs_0.9.1
## [25] splines_3.2.3     lme4_1.1-14    stringr_1.2.0
## [28] htmlwidgets_0.9   igraph_1.1.0   munsell_0.4.3
## [31] shiny_1.0.5       httpuv_1.3.5   pkgconfig_2.0.1
## [34] base64enc_0.1-3   rstantools_1.3.0 htmltools_0.3.6
## [37] tidyselct_0.2.2   tibble_1.3.4   codetools_0.2-15
## [40] threejs_0.3.1     matrixStats_0.52.2 MASS_7.3-47
## [43] grid_3.2.3        nlme_3.1-131   xtable_1.8-2
## [46] gtable_0.2.0      magrittr_1.5    scales_0.5.0
## [49] stringi_1.1.5     reshape2_1.4.2 dygraphs_1.1.1.4
## [52] xts_0.10-0        tools_3.2.3     glue_1.2.0
## [55] markdown_0.8      shinystan_2.4.0 purrr_0.2.4
## [58] crosstalk_1.0.0   rsconnect_0.8.5 parallel_3.2.3
## [61] yaml_2.1.14       inline_0.3.14   colorspace_1.3-2
## [64] knitr_1.17        bindr_0.1

```

## Appendix: Licenses

- Code © 2017, Aki Vehtari, licensed under BSD-3.
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